



Single-Dish Radio Telescopes

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Telescope Optics

Prime Focus: Retractable boom

Gregorian Focus: 8-m subreflector - 6-degrees of freedom



Telescope Structure

- *Fully Steerable*
 - *Elevation Limit: 5°*
 - *Can observe 85% of the entire Celestial Sphere*
- *Slew Rates: Azimuth - 40°/min; Elevation - 20°/min*



Telescope Structure

Blind Pointing:
(1 point/focus)

$$\sigma_2 \approx 5 \text{ arcsec}$$

$$\sigma(\text{focus}) \approx 2.5 \text{ mm}$$

Offset Pointing:
(90 min)

$$\sigma_2 \approx 2.7 \text{ arcsec}$$

$$\sigma(\text{focus}) \approx 1.5 \text{ mm}$$

Continuous Tracking:
(30 min)

$$\sigma_2 \approx 1 \text{ arcsec}$$

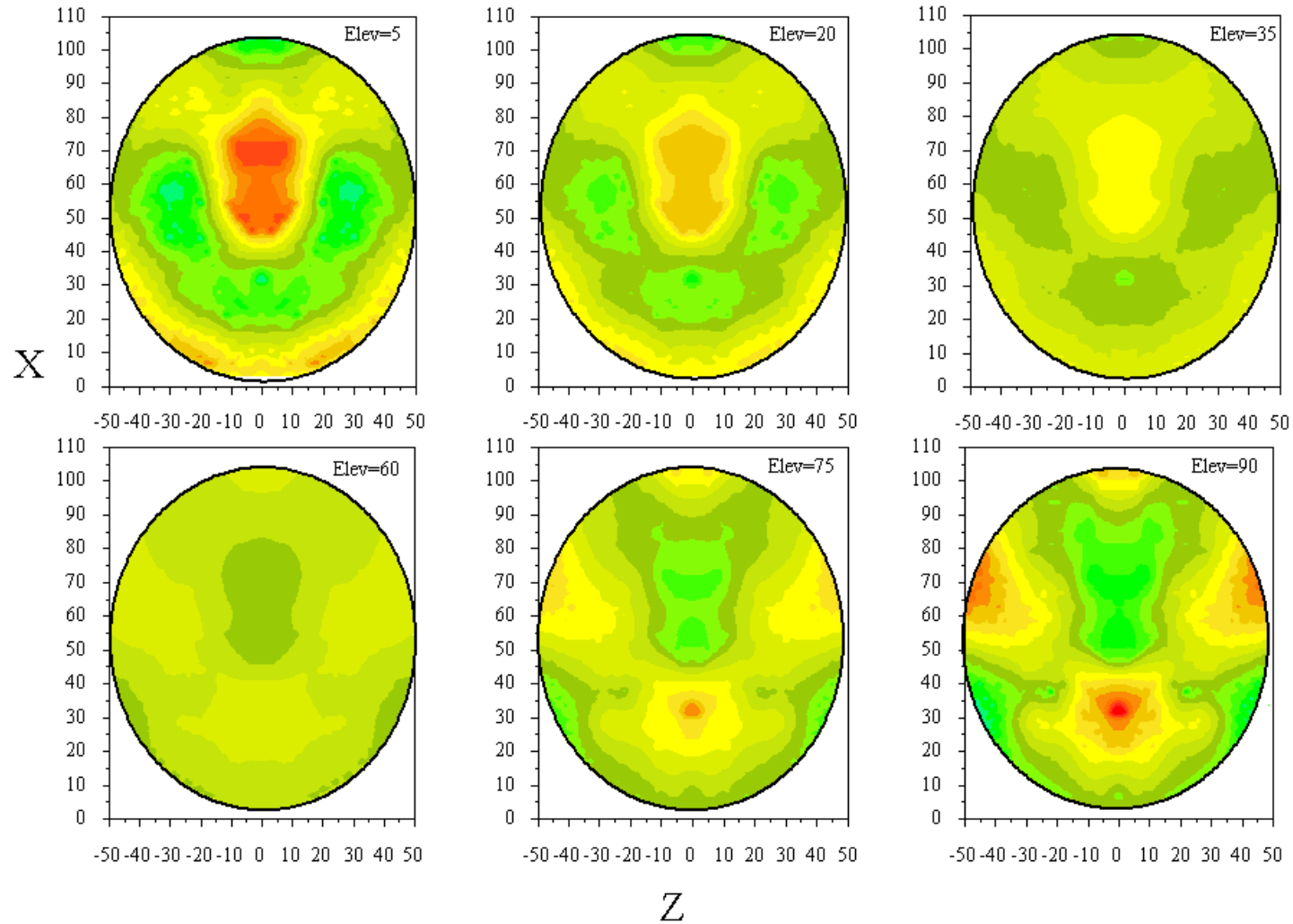


Telescope Structure

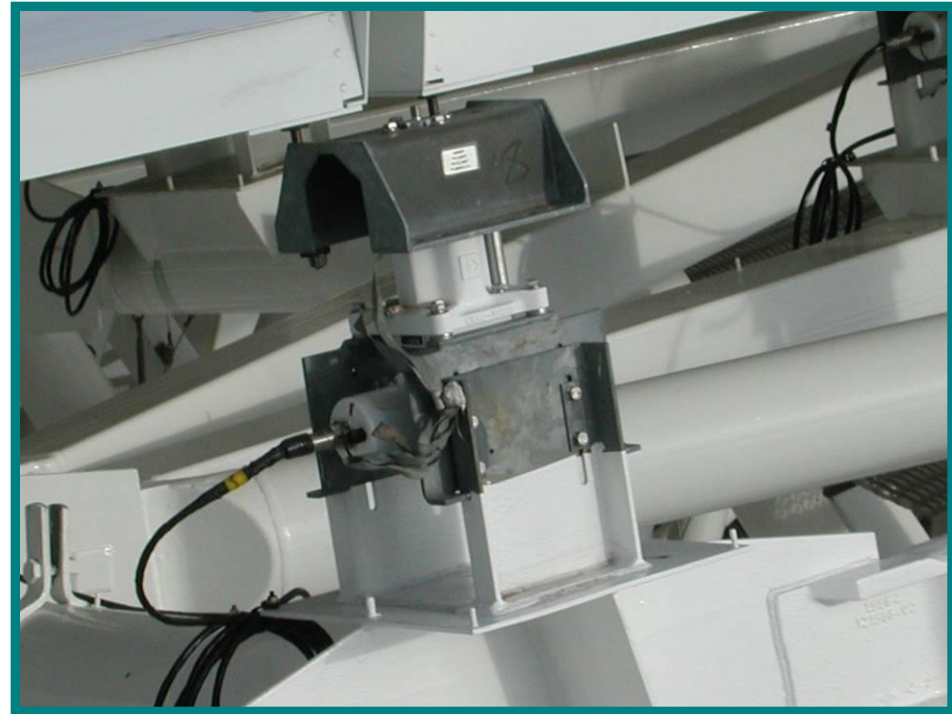


Active Surface

Surface Deformations from Finite Element Model



Active Surface



Telescope Optics

Rotating Turret with 8 receiver bays

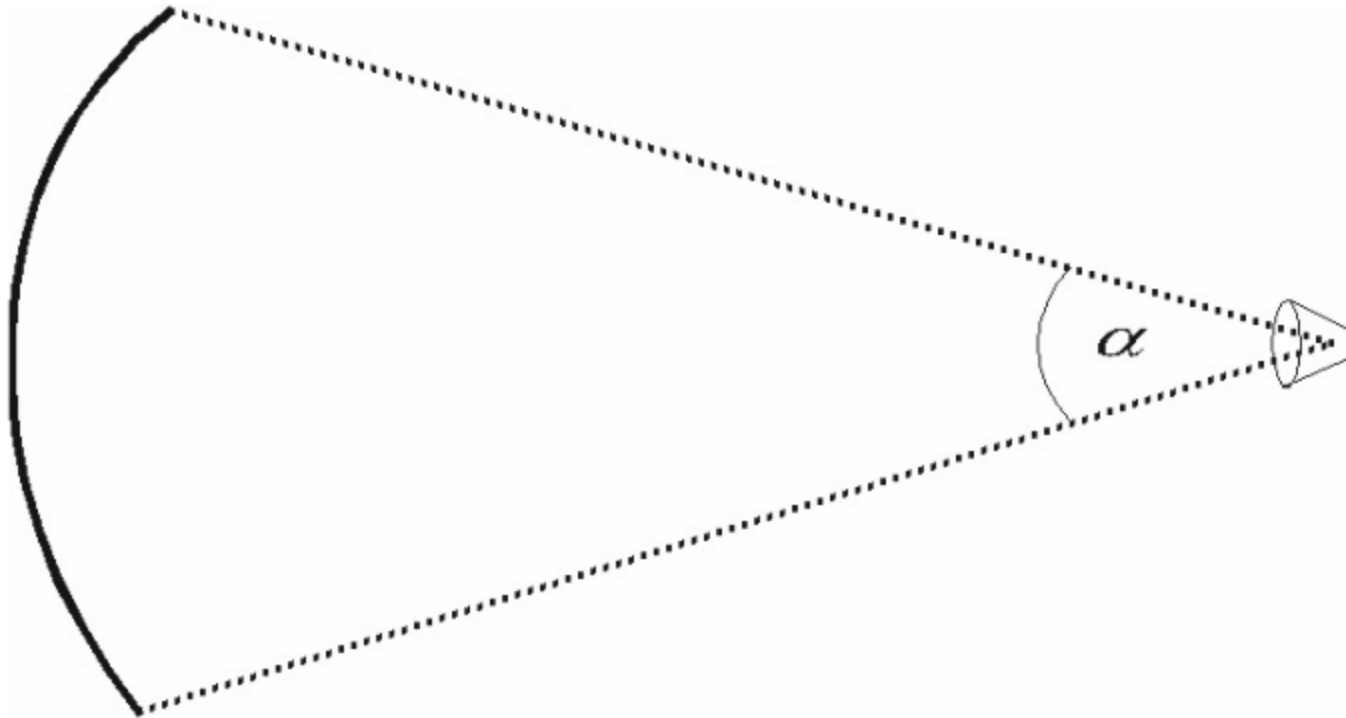


Receivers

Receiver	Operating Range	Status
Prime Focus 1	0.29—0.92 GHz	Commissioned
Prime Focus 2	0.910—1.23 GHz	Commissioned
L Band	1.15—1.73 GHz	Commissioned
S Band	1.73—2.60 GHz	Commissioned
C Band	3.95—5.85 GHz	Being Upgraded
X Band	8.2—10.0 GHz	Commissioned
Ku Band	12.4—15.4 GHz	Commissioned
K Band 7-pixel	18—26.5 GHz	Commissioned
Ka Band	26—40 GHz	Commissioned
Q Band	40—50 GHz	Commissioned
W Band	68—92 GHz	Commissioned
Penn Array	86—94 GHz	Being Upgraded



Reflector Feeds

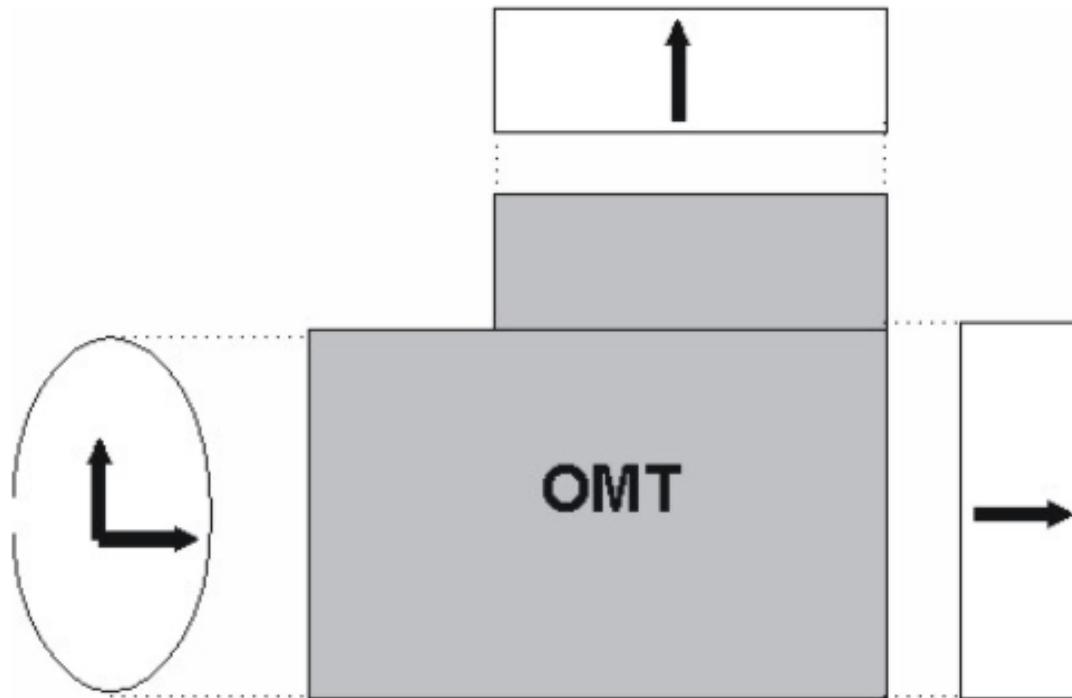




And More Feeds



Linear Polarization



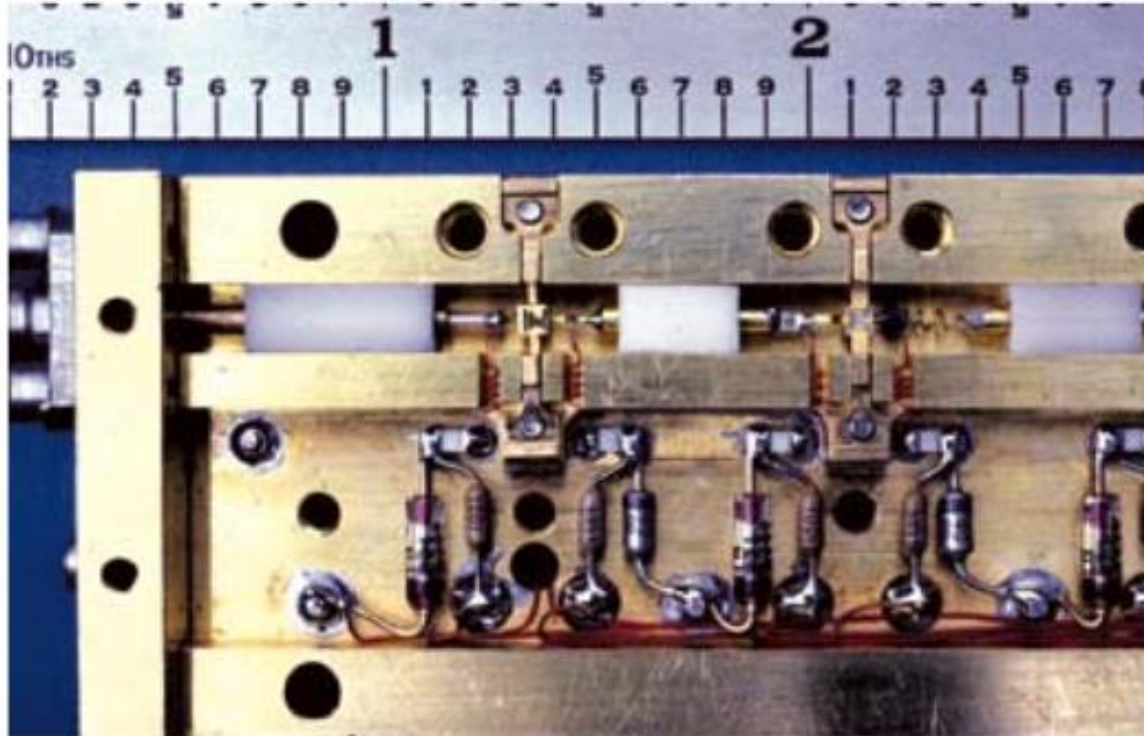
Orthomode Transducer

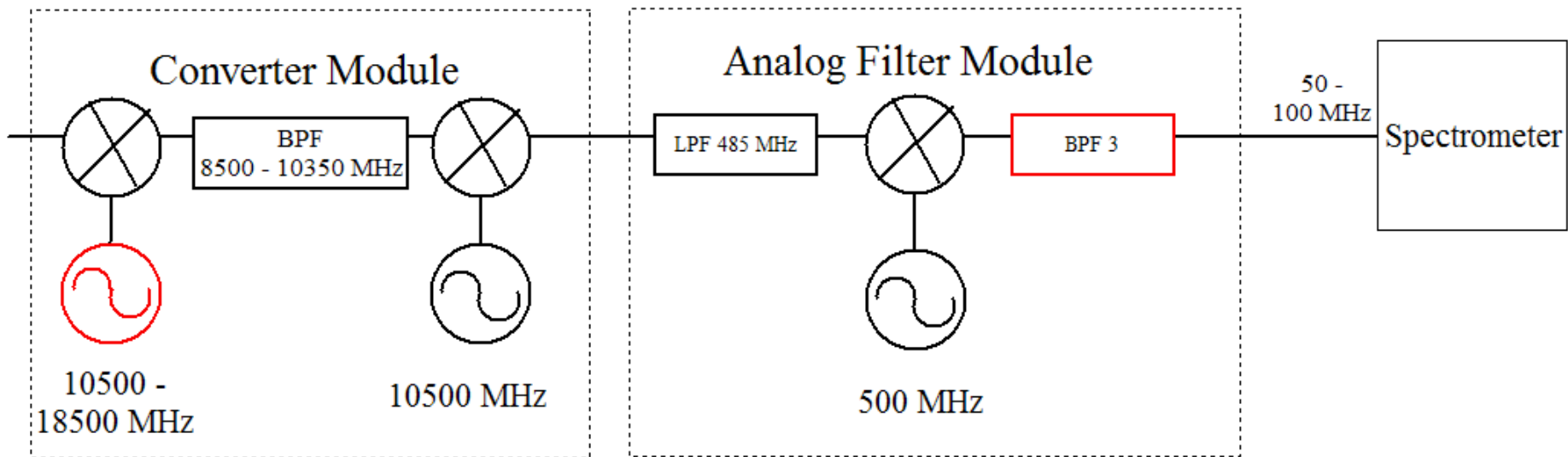
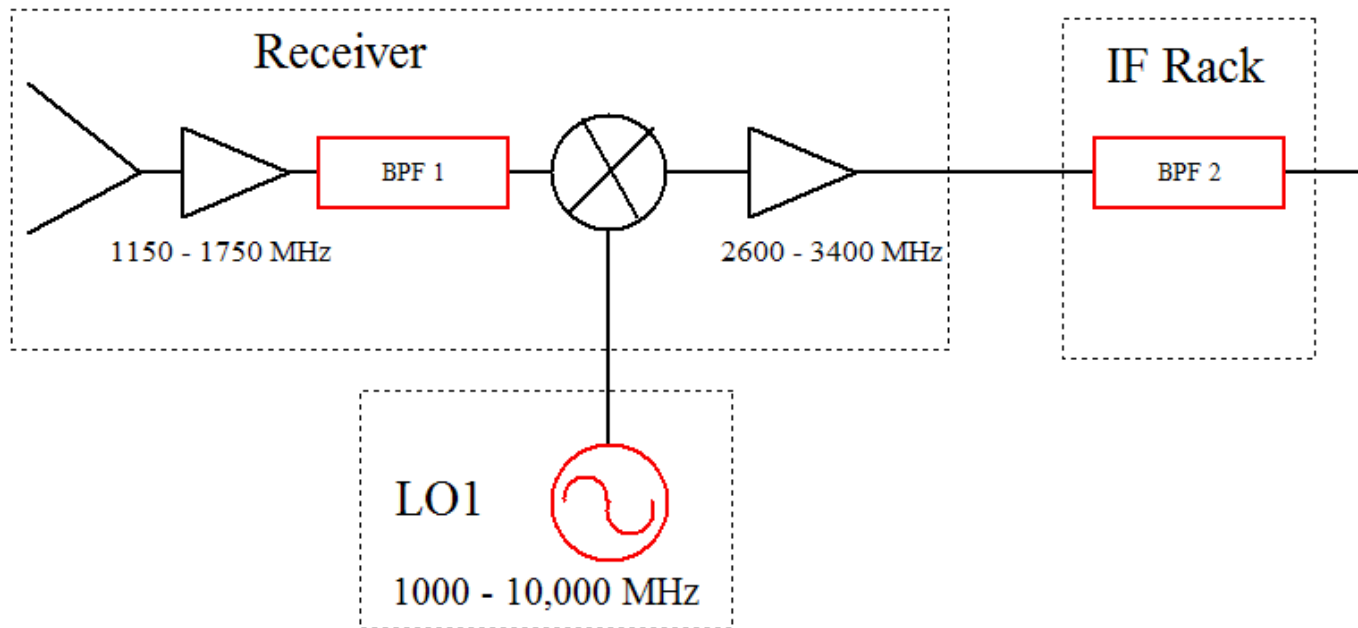


A Variety of OMTs



A HFET LNA





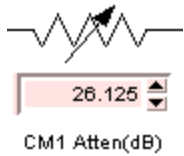
Typical Components



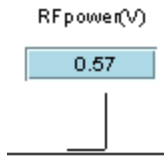
Amplifiers



Mixers



Attenuators



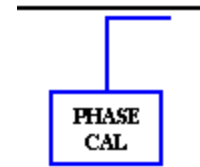
Power Detectors



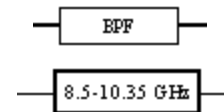
Synthesizers



Splitters



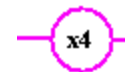
Couplers



Filters



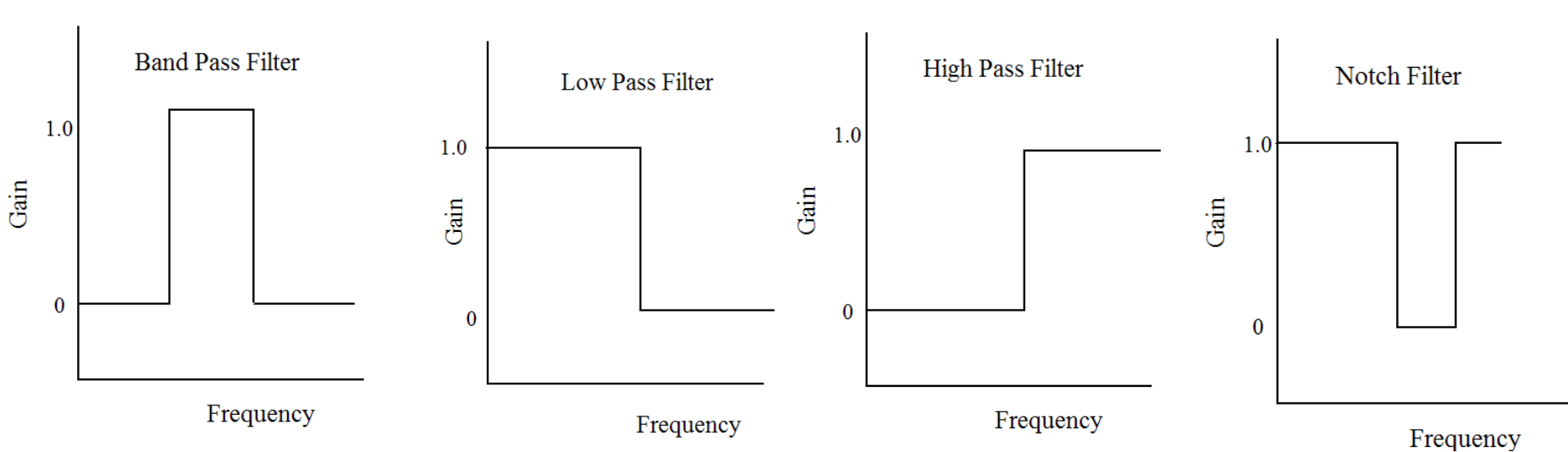
Switches



Multipliers



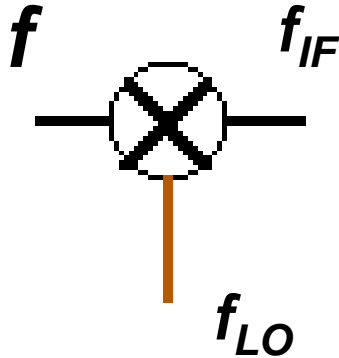
Types of Filters



Edges are smoother than illustrated



Types of Mixers

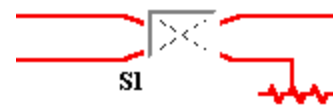
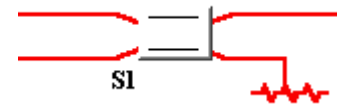
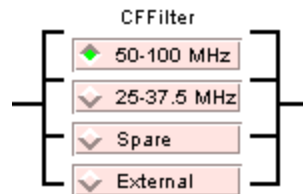
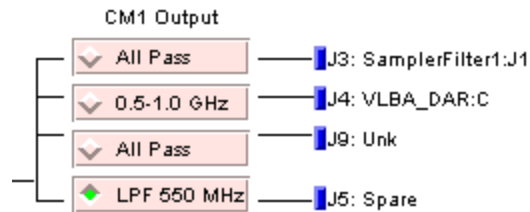
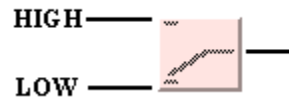


$$f_{IF} = n * f_{LO} + m * f$$

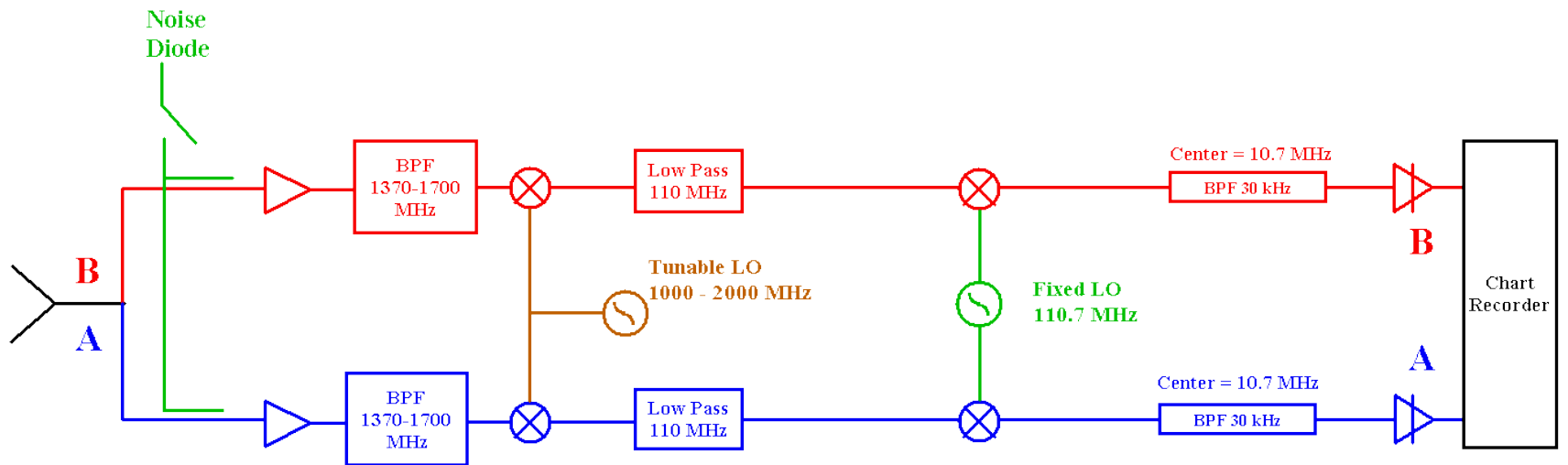
- ***n and m are positive or negative integers, usually 1 or -1***
- ***Up Conversion : $f_{IF} > f$***
- ***Down Conversion : $f_{IF} < f$***
- ***Lower Side Band : $f_{LO} > f$***
- Sense of frequency flips
- ***Upper Side Band : $f_{LO} < f$***



Example Switches



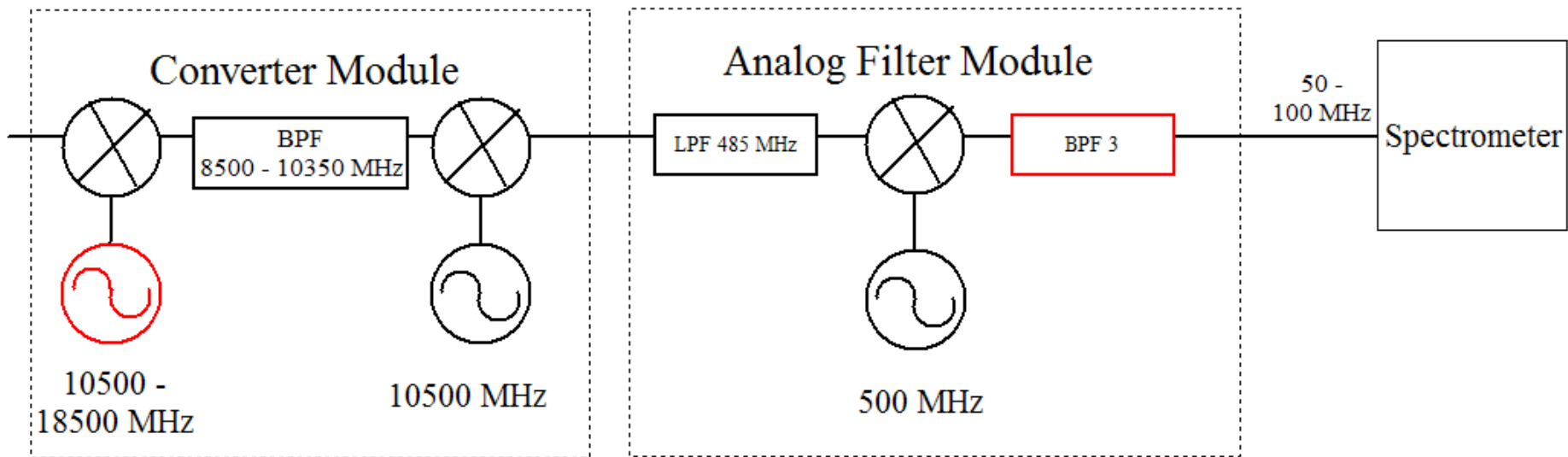
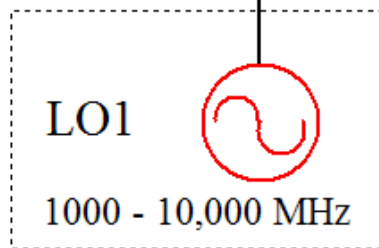
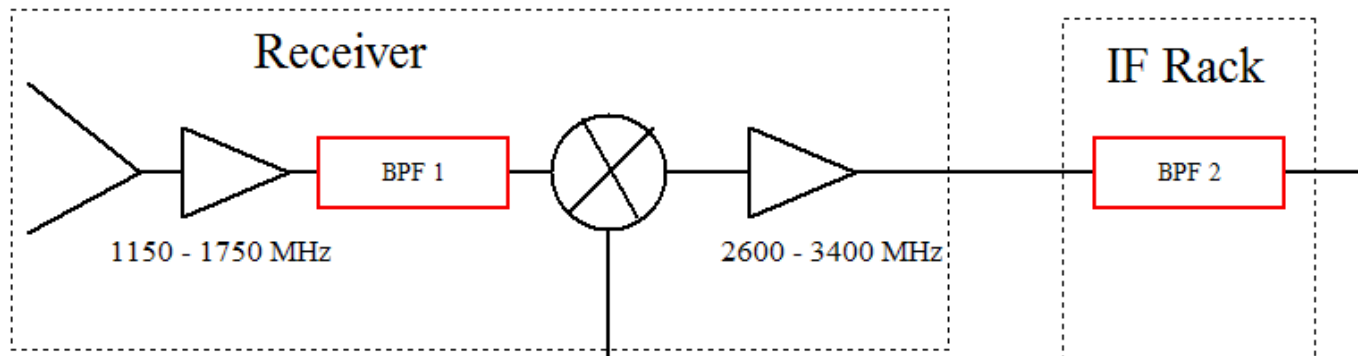
40-Ft System



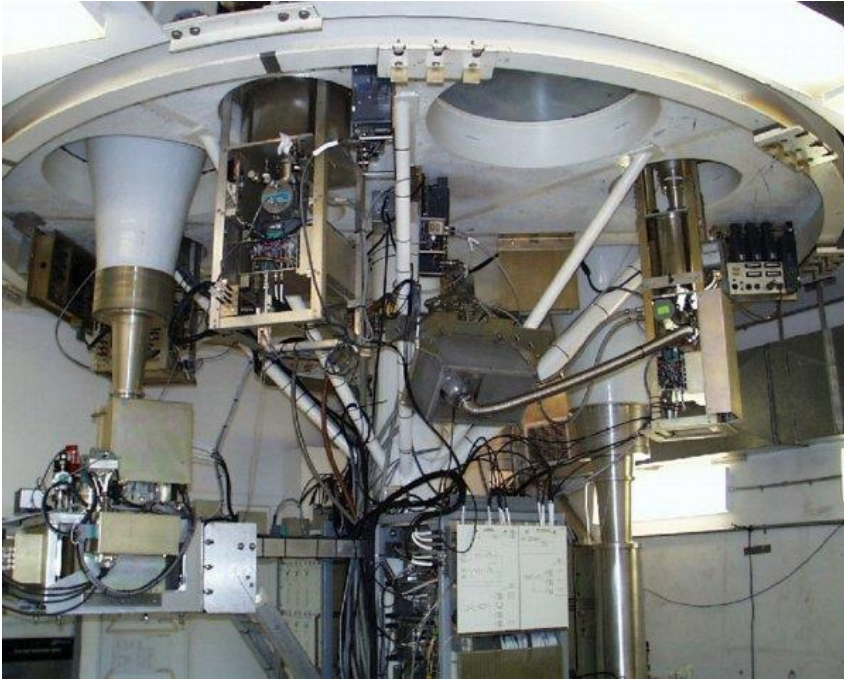
Quiz 1: Determine values for the first LO for the 40-ft when...

- Observing HI at 1420.41 MHz
- Observing OH at 1665.6 MHz

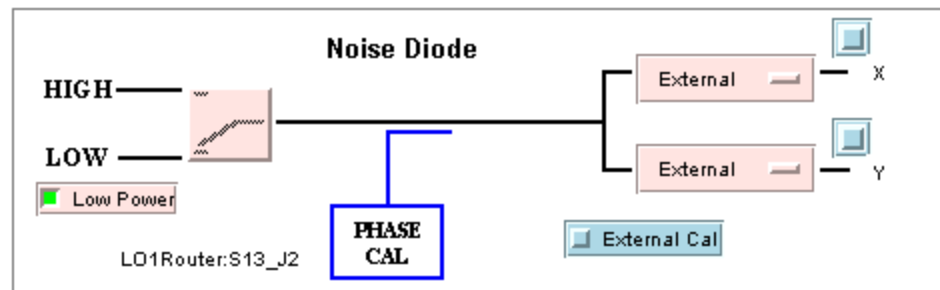
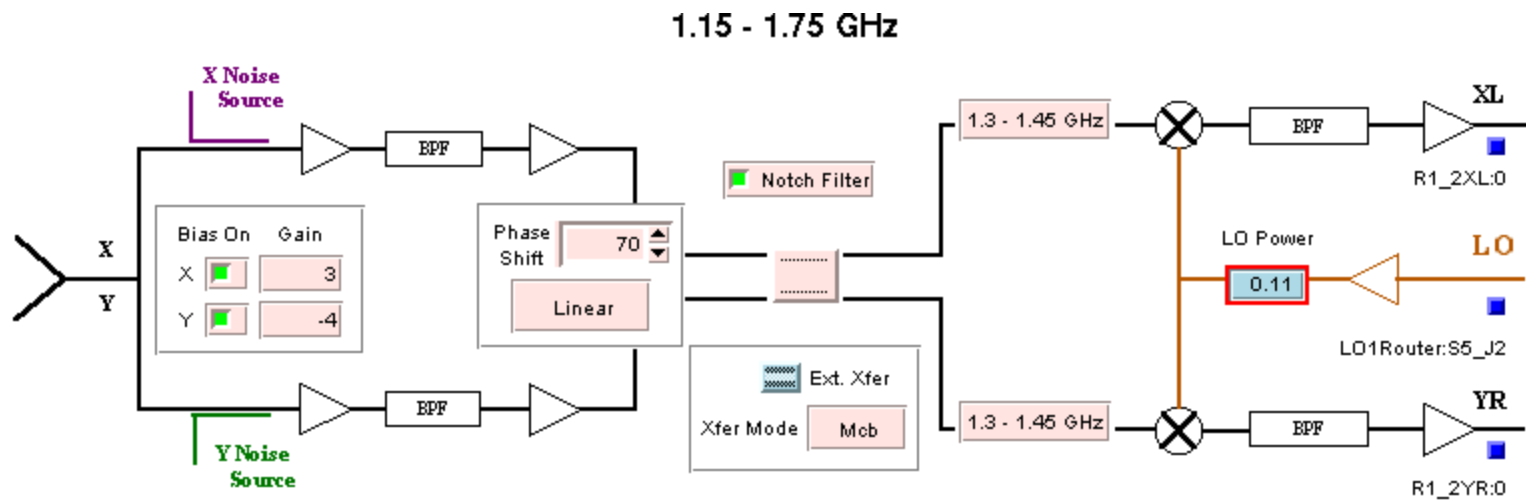




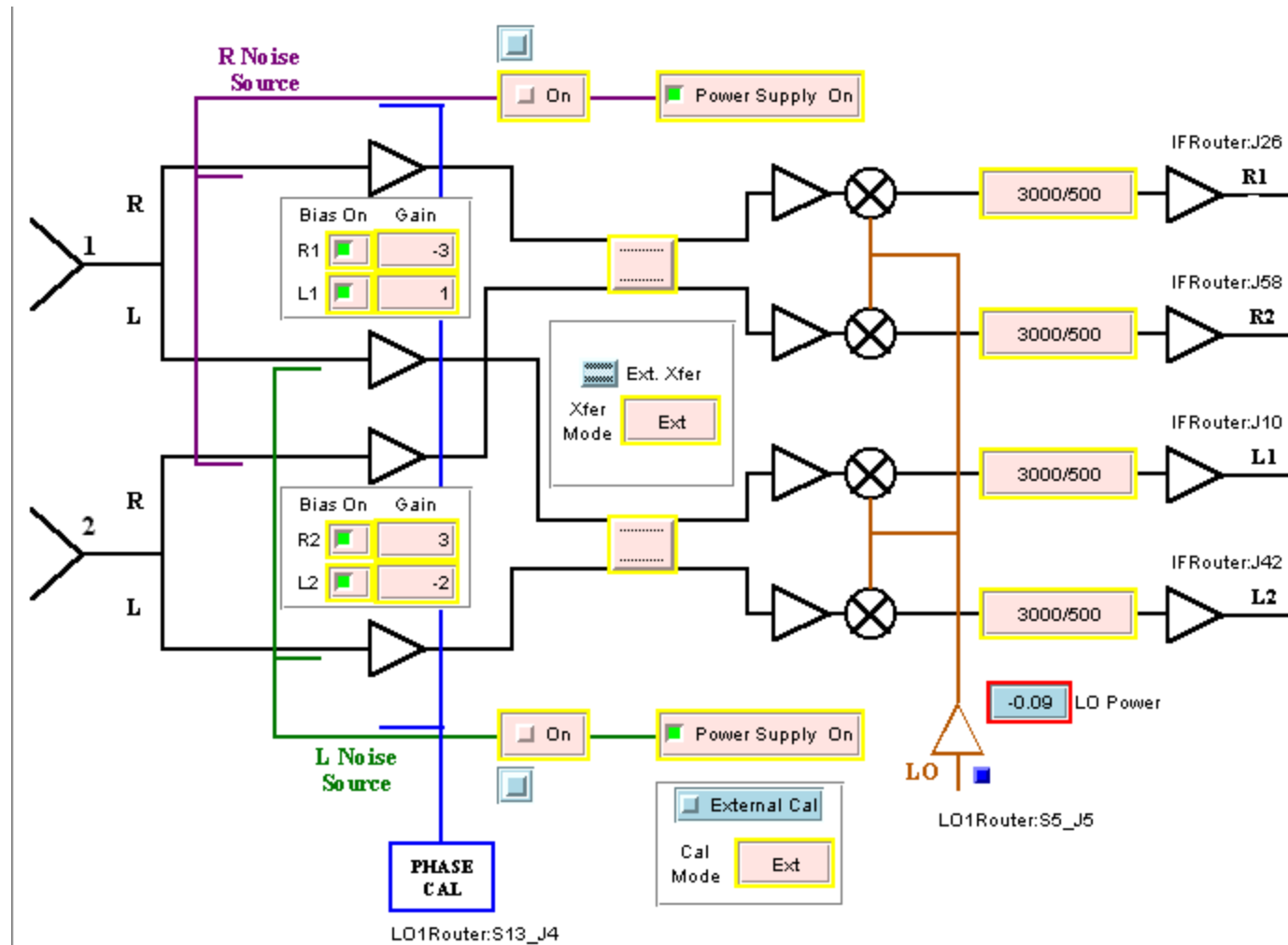
Receiver Room

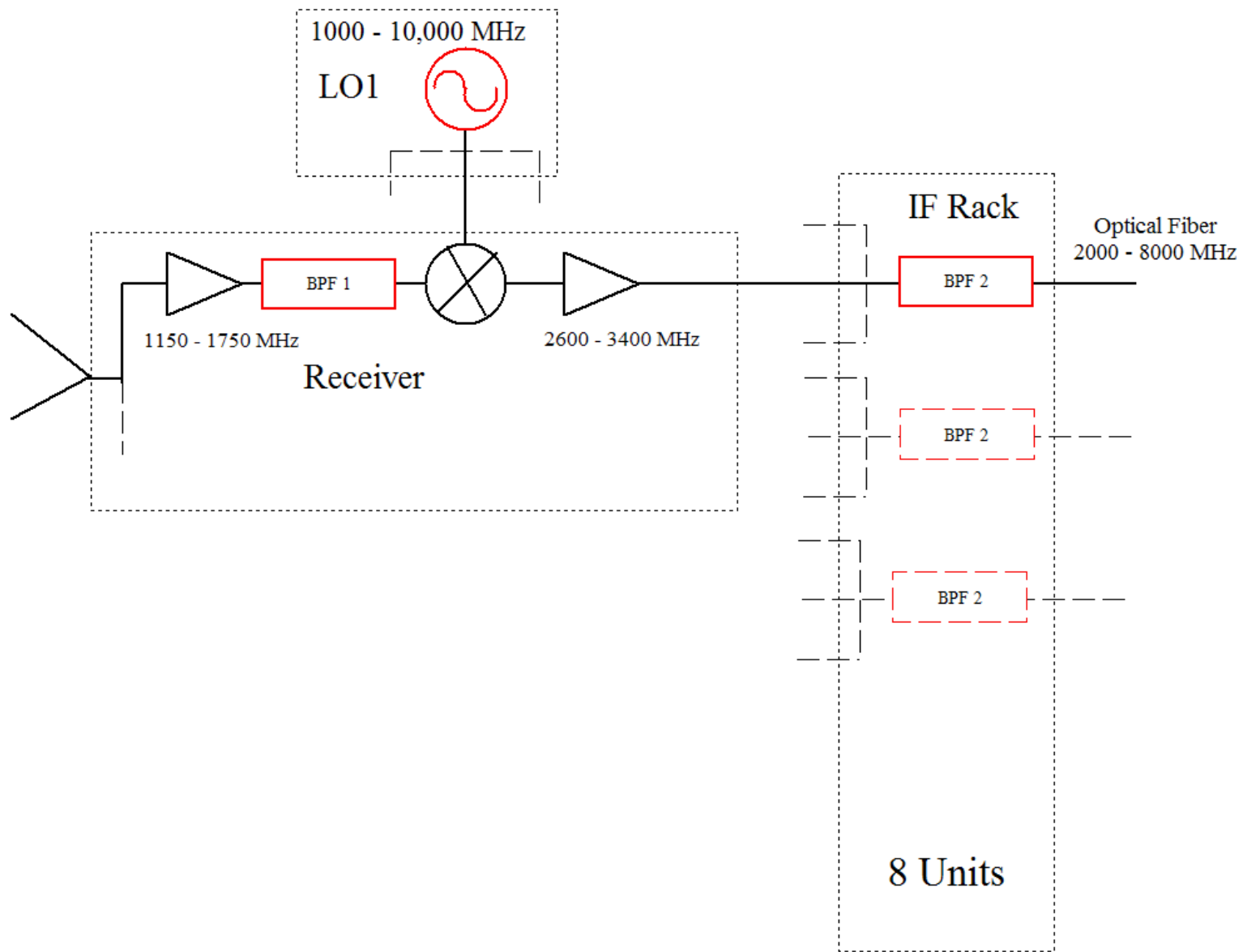


Typical Receiver

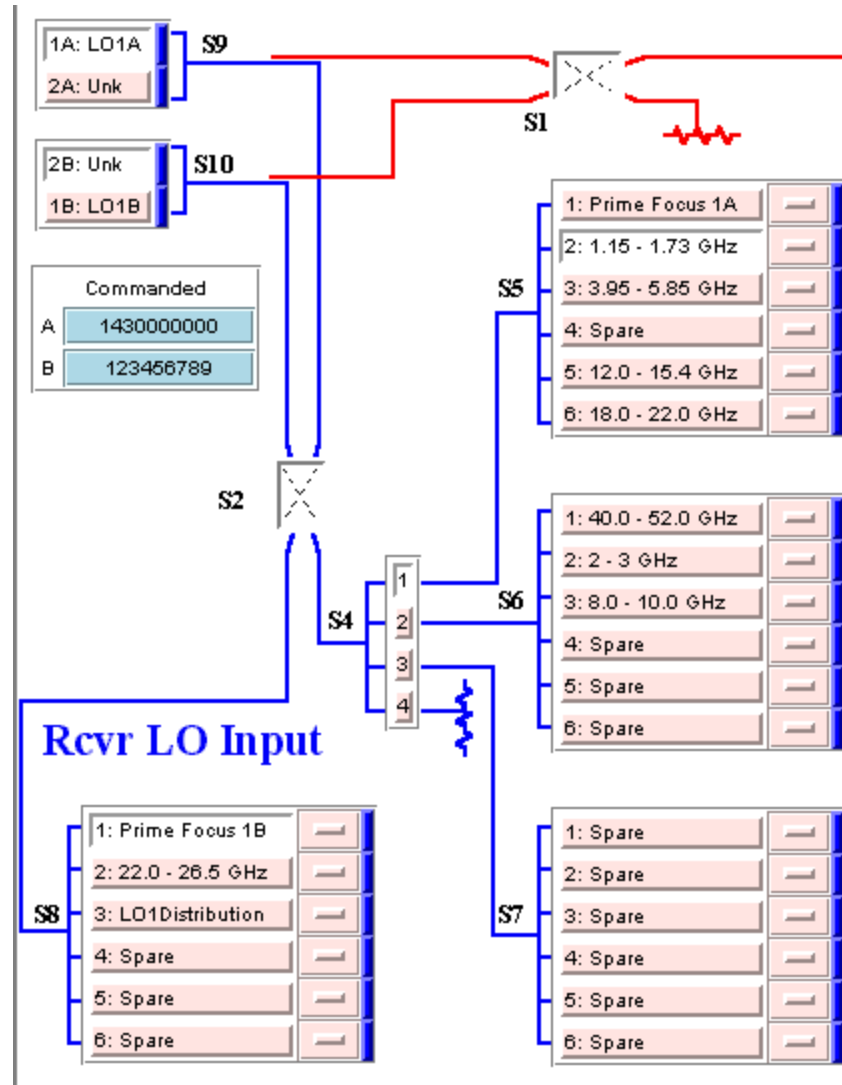


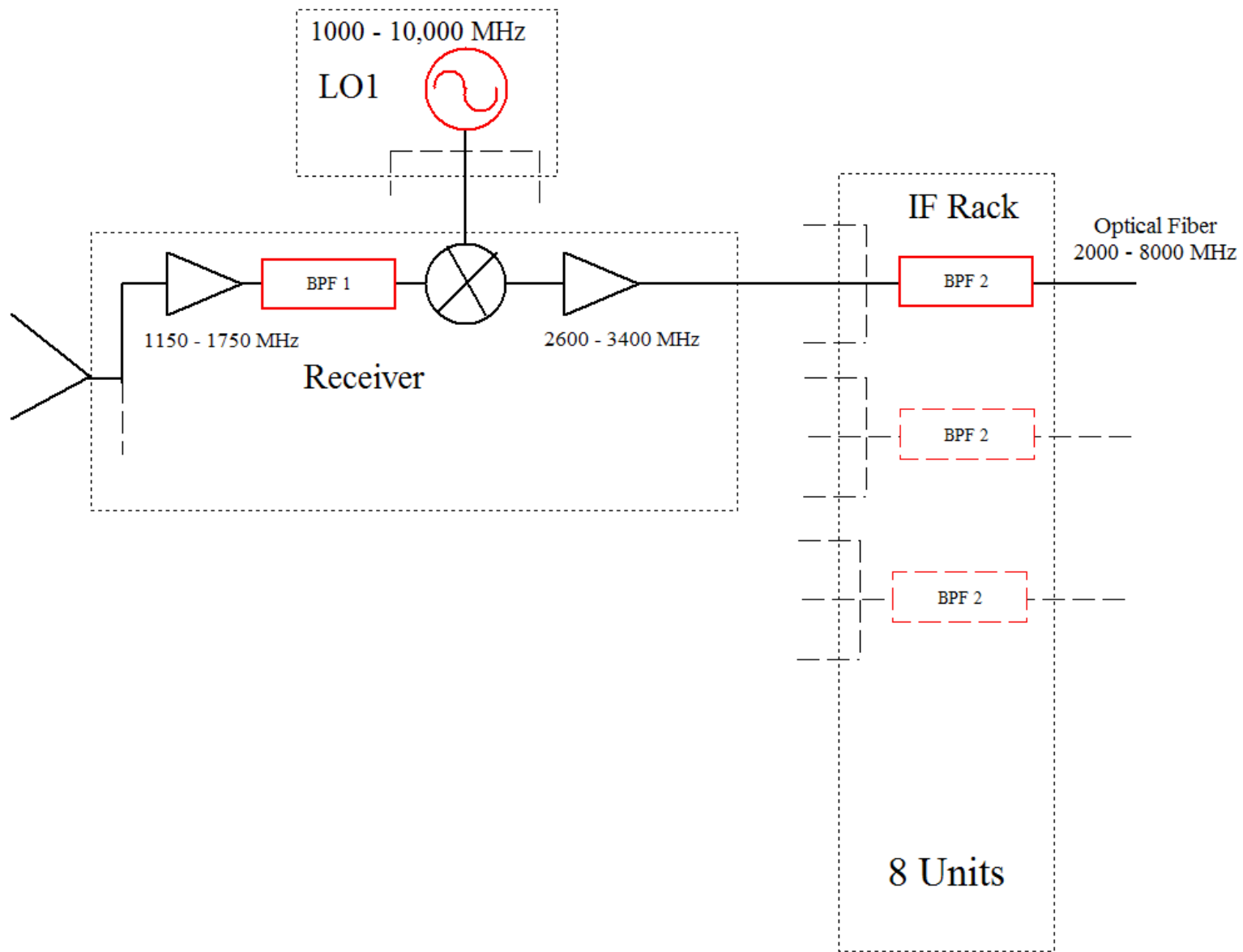
Multi-beam Receiver



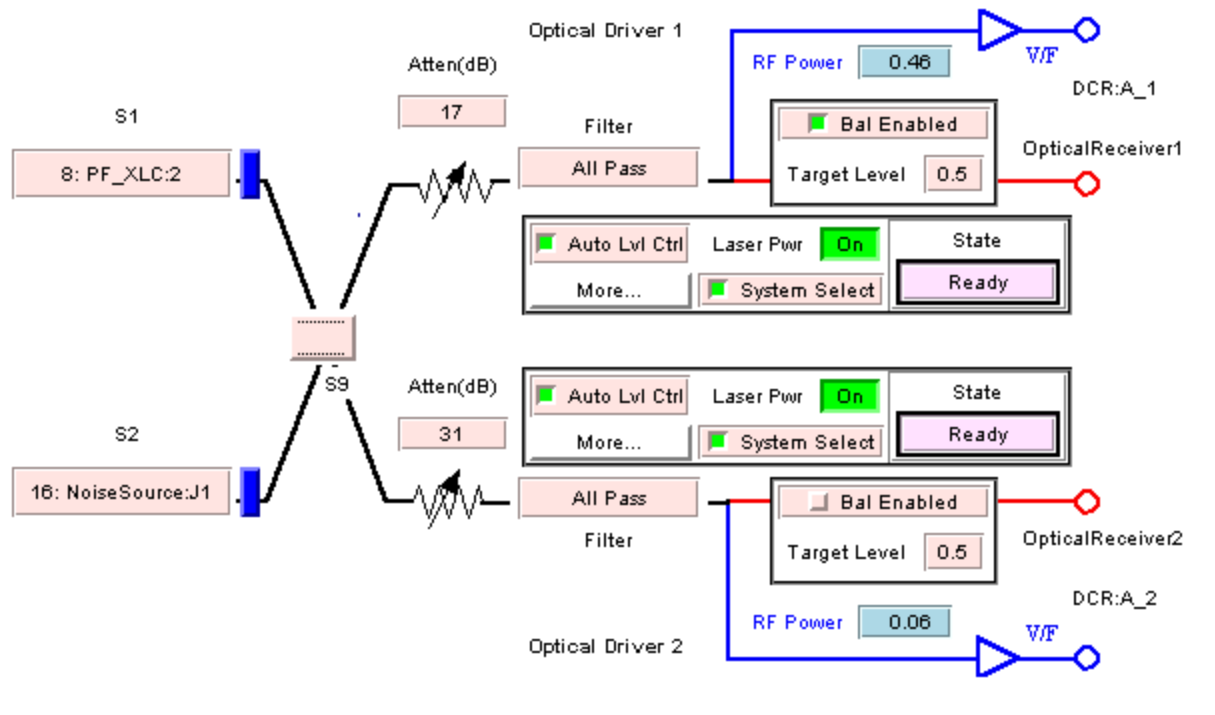


Local Oscillator and Switching Matrix

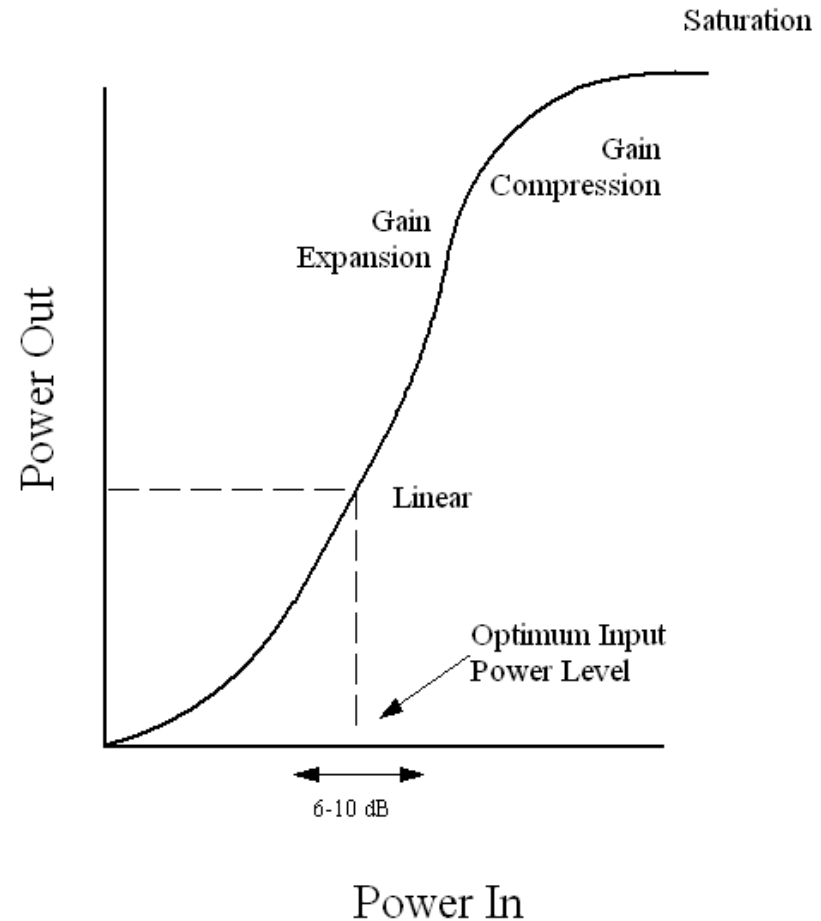
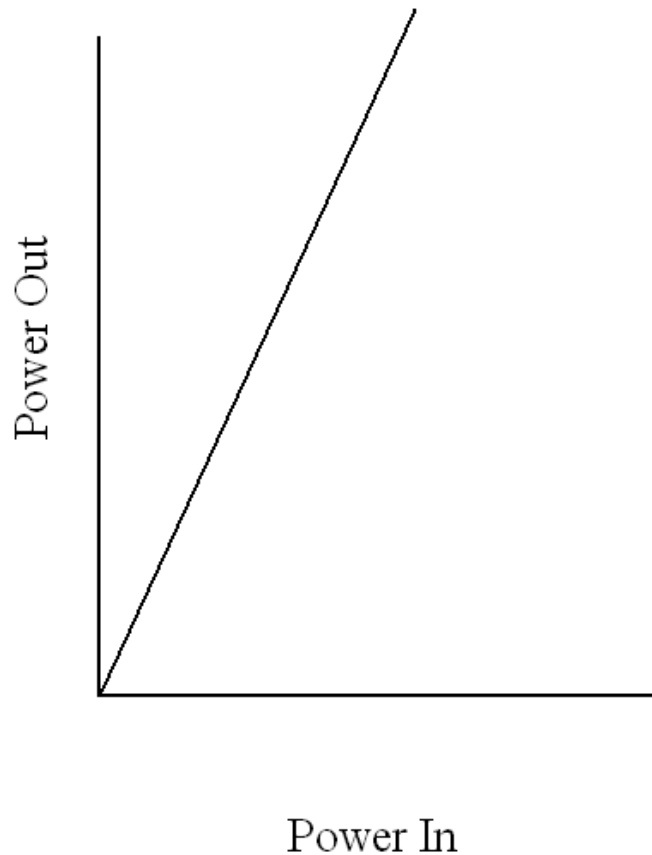


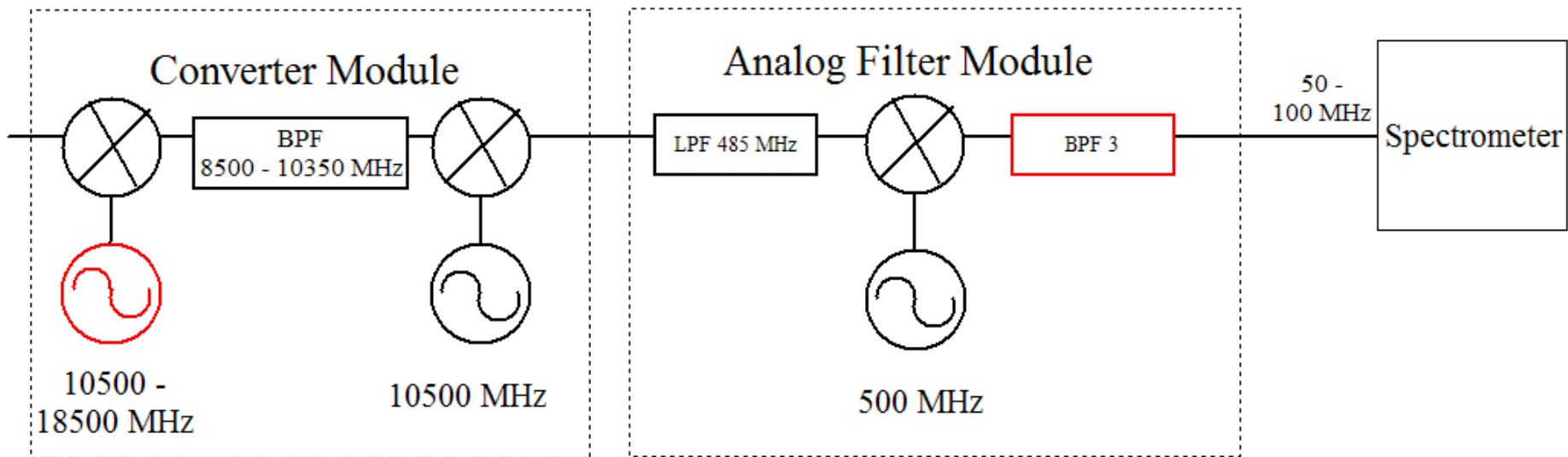
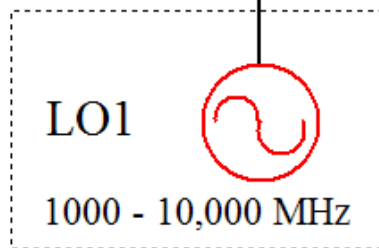
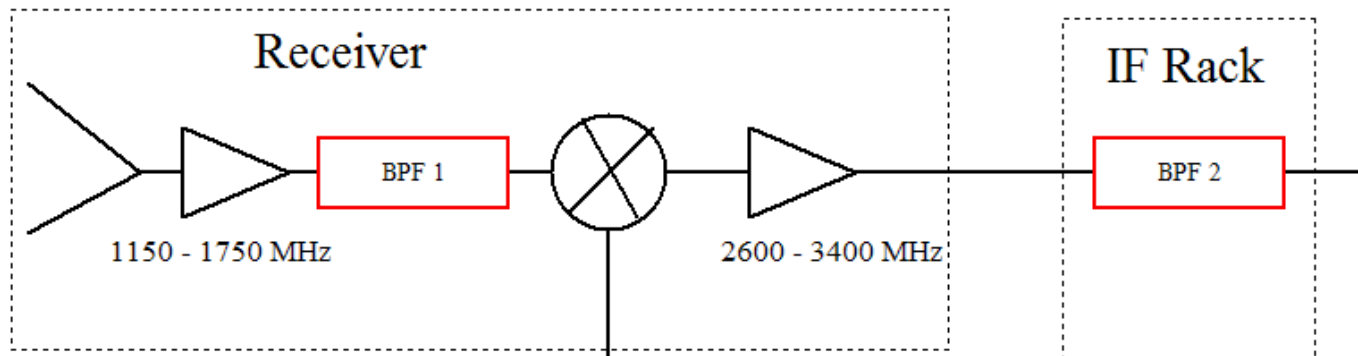


IF Rack – Input switching Matrix, IF Filters, Power Balancing Attenuators, and Drivers for 8 Optical Fibers

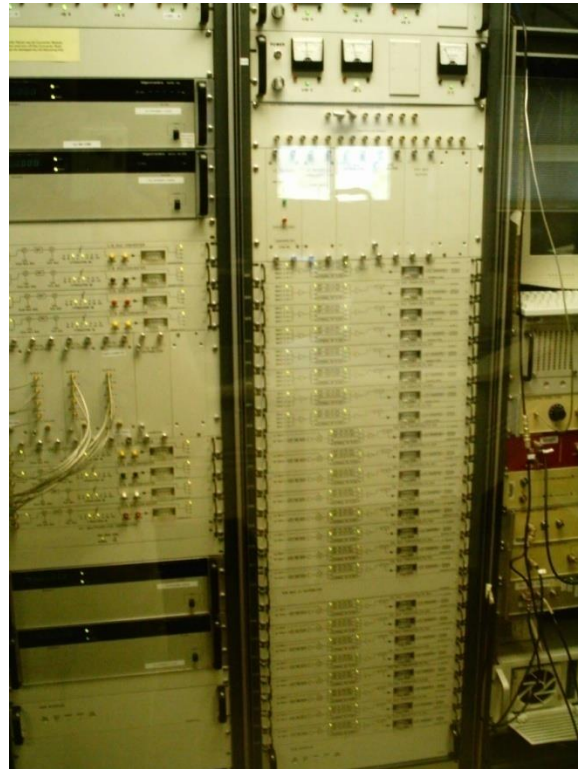


Power Balancing/Leveling and Non-Linearity

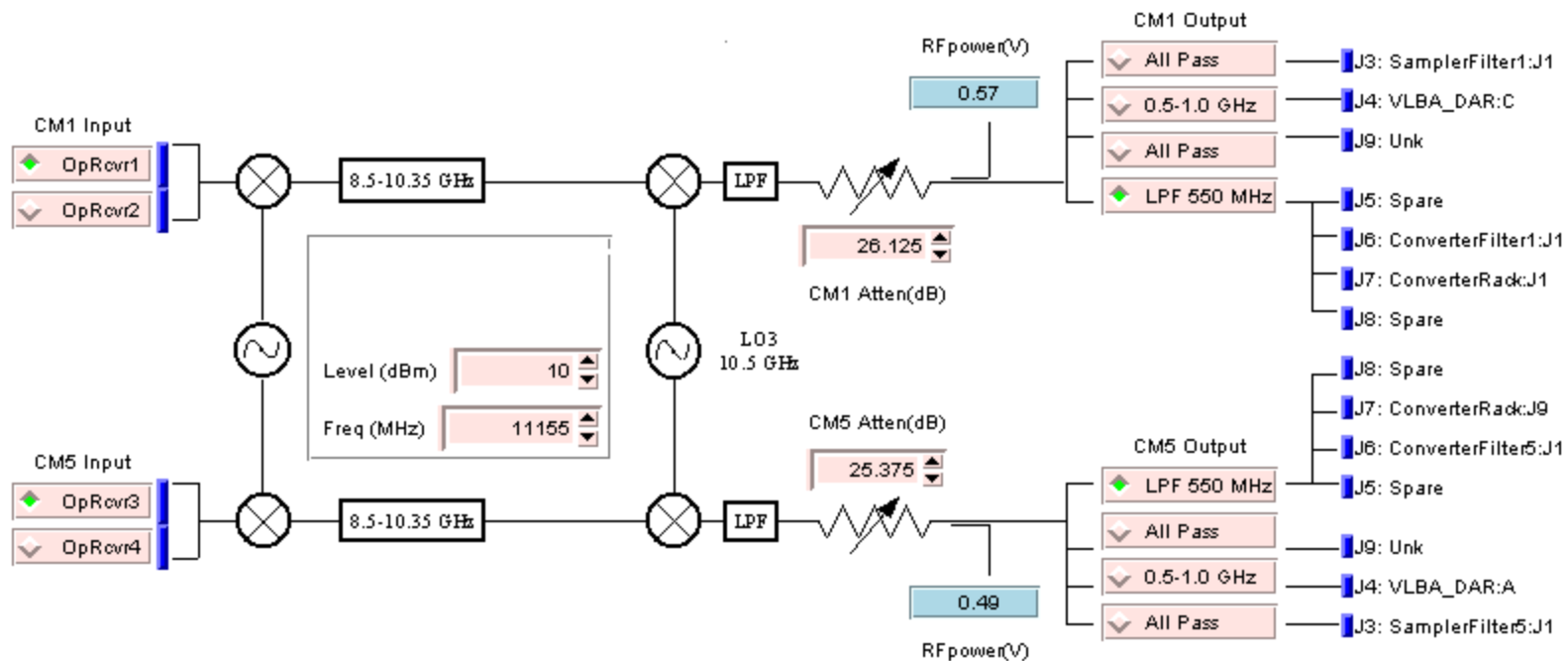




Converter and Analog Filter Racks, Spectrometer

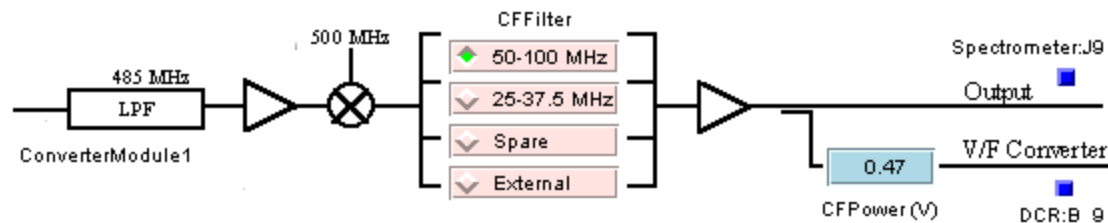


Converter Rack – Receivers for Optical Fibers, LO2 and LO3, Power Balancing Attenuators, Output Switches to Backends and AFR

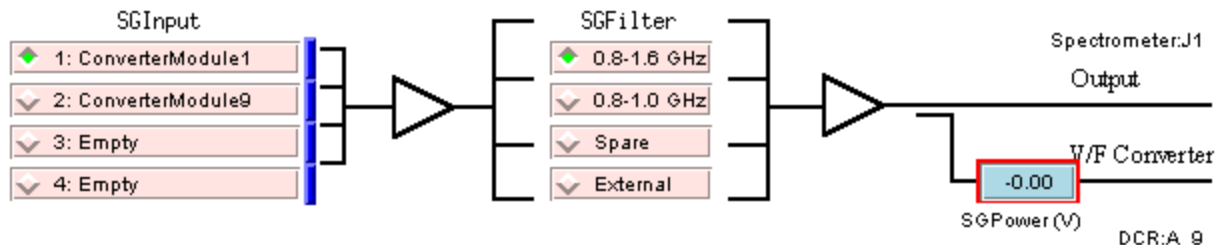


Analog Filter Rack

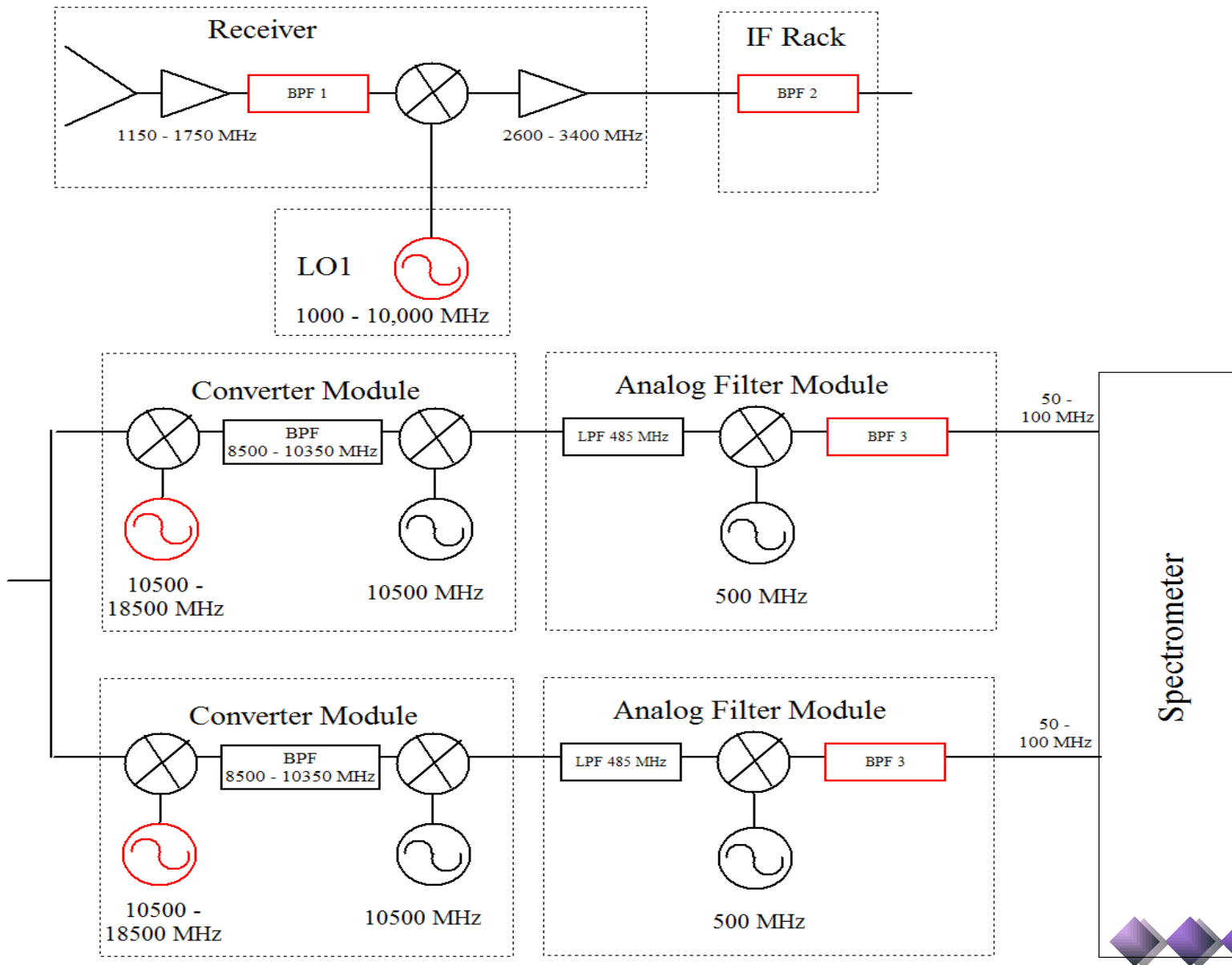
**For 12.5 and 50 MHz Slow-Speed Spectrometer Samplers :
LO4 and Filters**



**For 200 and 800 MHz High-Speed Spectrometer Samplers :
Input Switches and Filters.**



Quiz 2: Determine values for red components

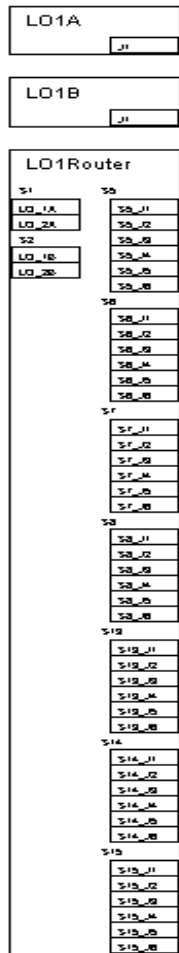


Quiz 2: Determine values for red components

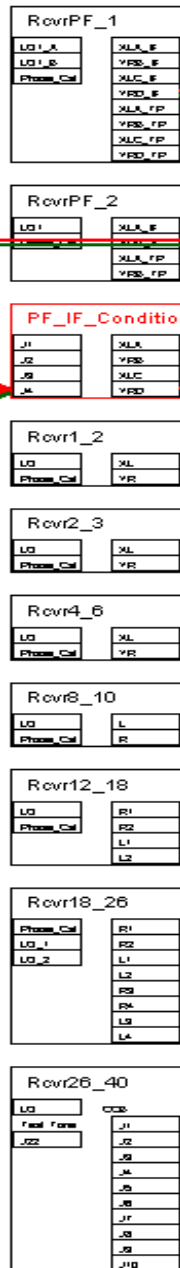
- Goal : Observe simultaneously 1420 MHz and 1665 MHz with the 50 MHz wide (75 MHz center frequency) mode of the Spectrometer
- Parameters:
 - BPF1 can be: 1100–1800, 1600-1750, 1300-1450, or 1100-1450 MHz
 - All mixers are LSB. Hint: first two mixers up convert, the last two down convert.
 - BPF2 can be : 2990-3010, 2960-3040, 2840-3160, 2360-3640, 5960-6040, 5840-6160, or 5360-6640 MHz
 - BPF3 can be : 50-100 or 25-37.5 MHz
 - See block diagram for other parameters
- Hint: Work from the receiver down the chain until you get stuck, then from Spectrometer up. Try 1420 MHz first, then add in 1665 MHz.
- Record values for LO1 and both LO2's; settings for BPF1, 2, and 3; and values for all Intermediate Frequencies.



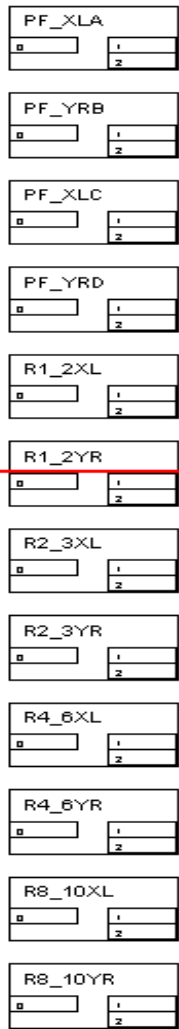
L.O. 1



RECEIVERS



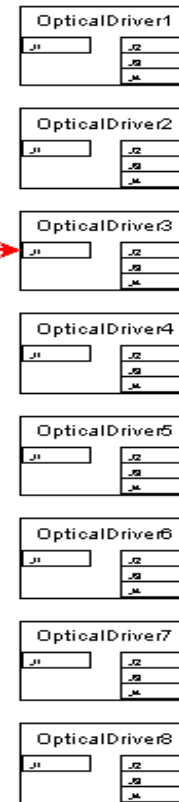
SPLITTERS



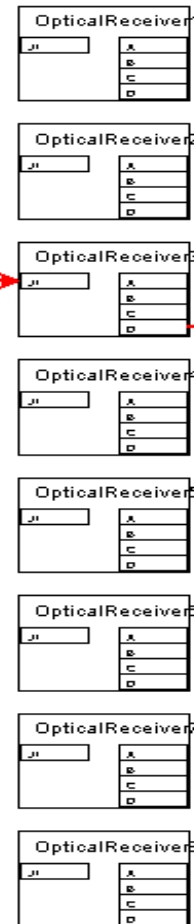
I.F. ROUTER

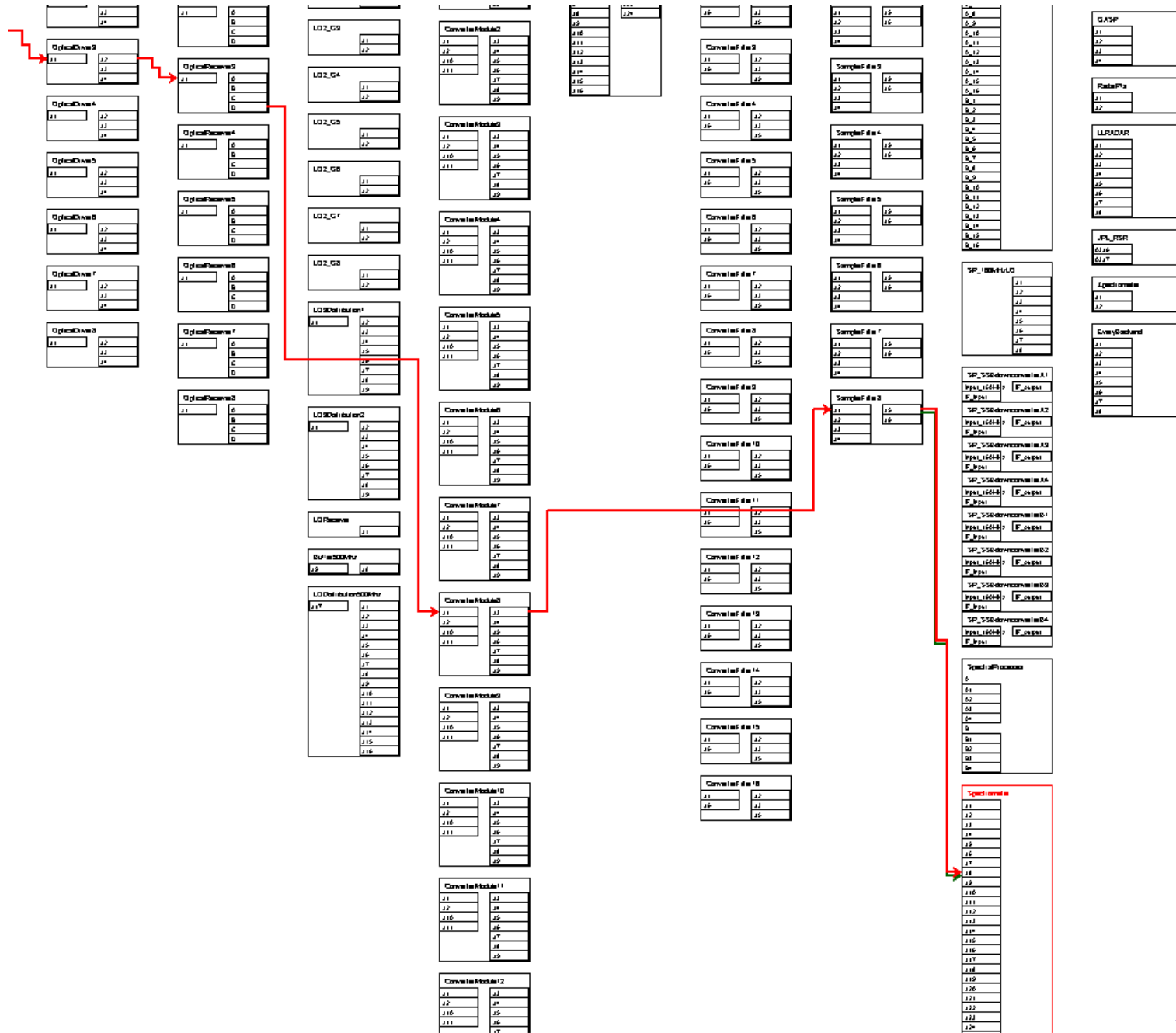


OPT. DRIVERS



OPT. RECEIVERS





Spectrum	
Sideband	lower
IF	1200
Sky	-2770
Bandwidth	0
Polarization	linear_y
Noise Diode	lowCal
Sinusoid	
IF	0
At LO	0
From: SamplerFilter8:J5	

```

Feed: RcvrPF_1:YRD_342
  Freq: 270 to 420 MHz
  Polarization: linear_y
  Horn: 1
Tone: RcvrPF_1:C342Y
  Freq: 0 MHz
Filter: RcvrPF_1:FL342_5Y
  Freq: 270 to 420 MHz
Mixer: RcvrPF_1:MXYRD
  LO: 1430 MHz
  Component -- LO1A:synthesizer
  Lower Sideband: IFo = 1430 - IFi
Filter: RcvrPF_1:FLYRD3
  Freq: 1040 to 1120 MHz
Filter: RcvrPF_1:FLYRD
  Freq: 960 to 1200 MHz
Attenuator: RcvrPF_1:ifChannelD
Output Port: RcvrPF_1:J4
Input Port: PF_IF_Conditioner:J4
Output Port: PF_IF_Conditioner:J8
Input Port: IFRouter:J23
Output Port: IFRouter:J67
Input Port: OpticalDriver3:J1
Attenuator: OpticalDriver3:attenuator
Output Port: OpticalDriver3:J2
Input Port: OpticalReceiver3:J1
Output Port: OpticalReceiver3:J5
Input Port: ConverterModule8:J1
Mixer: ConverterModule8:MX2
  LO: 13500 MHz
  Component -- LO2_G4:synthesizer
  Lower Sideband: IFo = 13500 - IFi
Filter: ConverterModule8:FL1
  Freq: 8500 to 10350 MHz
Mixer: ConverterModule8:MX3
  LO: 10500 MHz
  Component -- LO3Distribution1:synthesizer
  Lower Sideband: IFo = 10500 - IFi
Filter: ConverterModule8:FL2
  Freq: 0 to 2200 MHz

```

Spectrum	
Sideband	lower
IF	1200
Sky	-2770
Bandwidth	0
Polarization	linear_y
Noise Diode	lowCal
Sinusoid	
IF	0
At LO	0
From: SamplerFilter8:J5	

```

Mixer: RcvrPF_1:MXYRD
  LO: 1430 MHz
  Component -- LO1A:synthesizer
  Lower Sideband: IFo = 1430 - IFi
Filter: RcvrPF_1:FLYRD3
  Freq: 1040 to 1120 MHz
Filter: RcvrPF_1:FLYRD
  Freq: 960 to 1200 MHz
Attenuator: RcvrPF_1:ifChannelD
Output Port: RcvrPF_1:J4
Input Port: PF_IF_Conditioner:J4
Output Port: PF_IF_Conditioner:J8
Input Port: IFRouter:J23
Output Port: IFRouter:J67
Input Port: OpticalDriver3:J1
Attenuator: OpticalDriver3:attenuator
Output Port: OpticalDriver3:J2
Input Port: OpticalReceiver3:J1
Output Port: OpticalReceiver3:J5
Input Port: ConverterModule8:J1
Mixer: ConverterModule8:MX2
  LO: 13500 MHz
  Component -- LO2_G4:synthesizer
  Lower Sideband: IFo = 13500 - IFi
Filter: ConverterModule8:FL1
  Freq: 8500 to 10350 MHz
Mixer: ConverterModule8:MX3
  LO: 10500 MHz
  Component -- LO3Distribution1:synthesizer
  Lower Sideband: IFo = 10500 - IFi
Filter: ConverterModule8:FL2
  Freq: 0 to 2200 MHz
Attenuator: ConverterModule8:AT1
Output Port: ConverterModule8:J3
Input Port: SamplerFilter8:J1
Filter: SamplerFilter8:FL1
  Freq: 800 to 1600 MHz
Output Port: SamplerFilter8:J5
Input Port: Spectrometer:J8

```

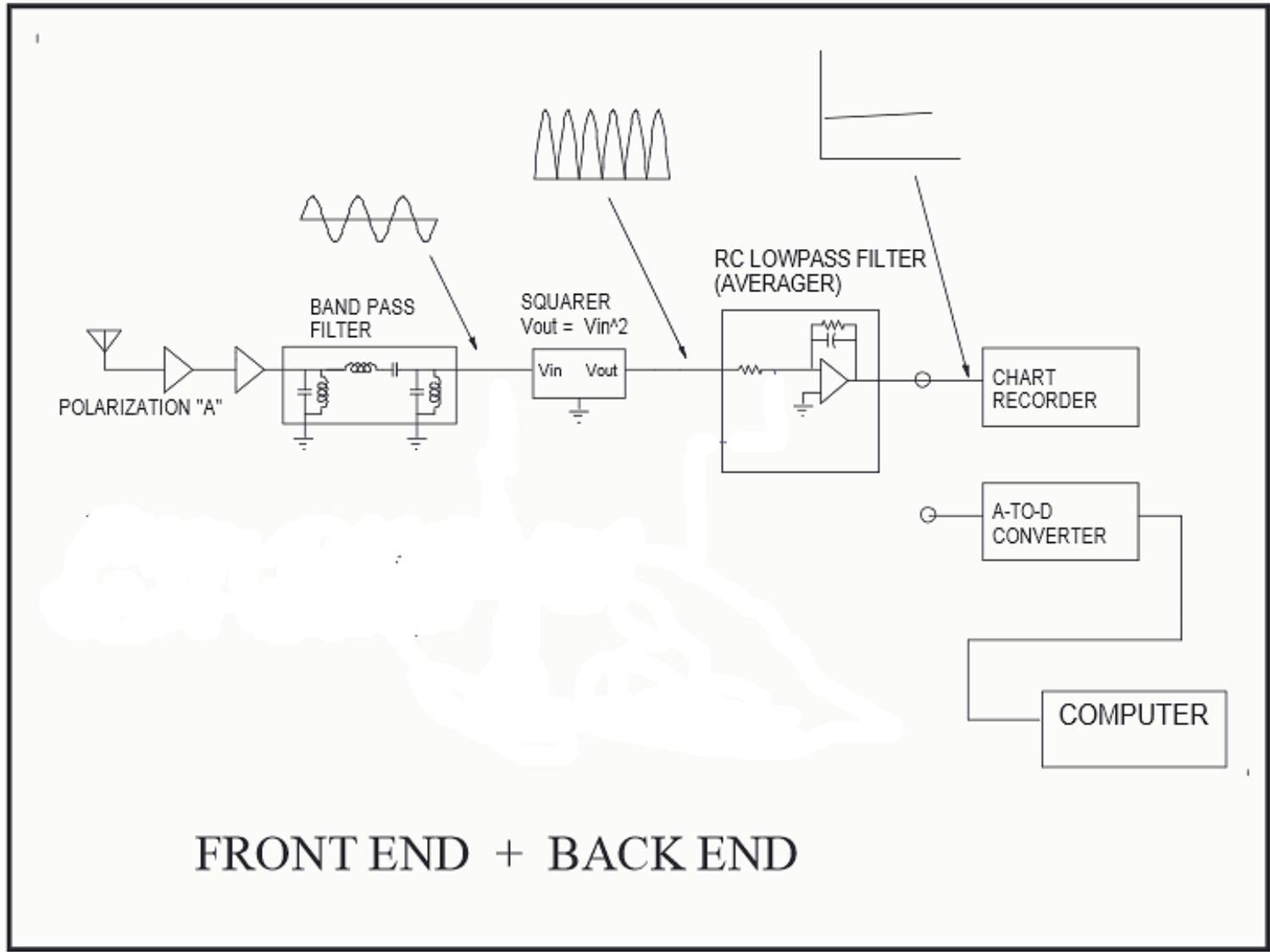


GBT – Astrid does all the hard work for you.....

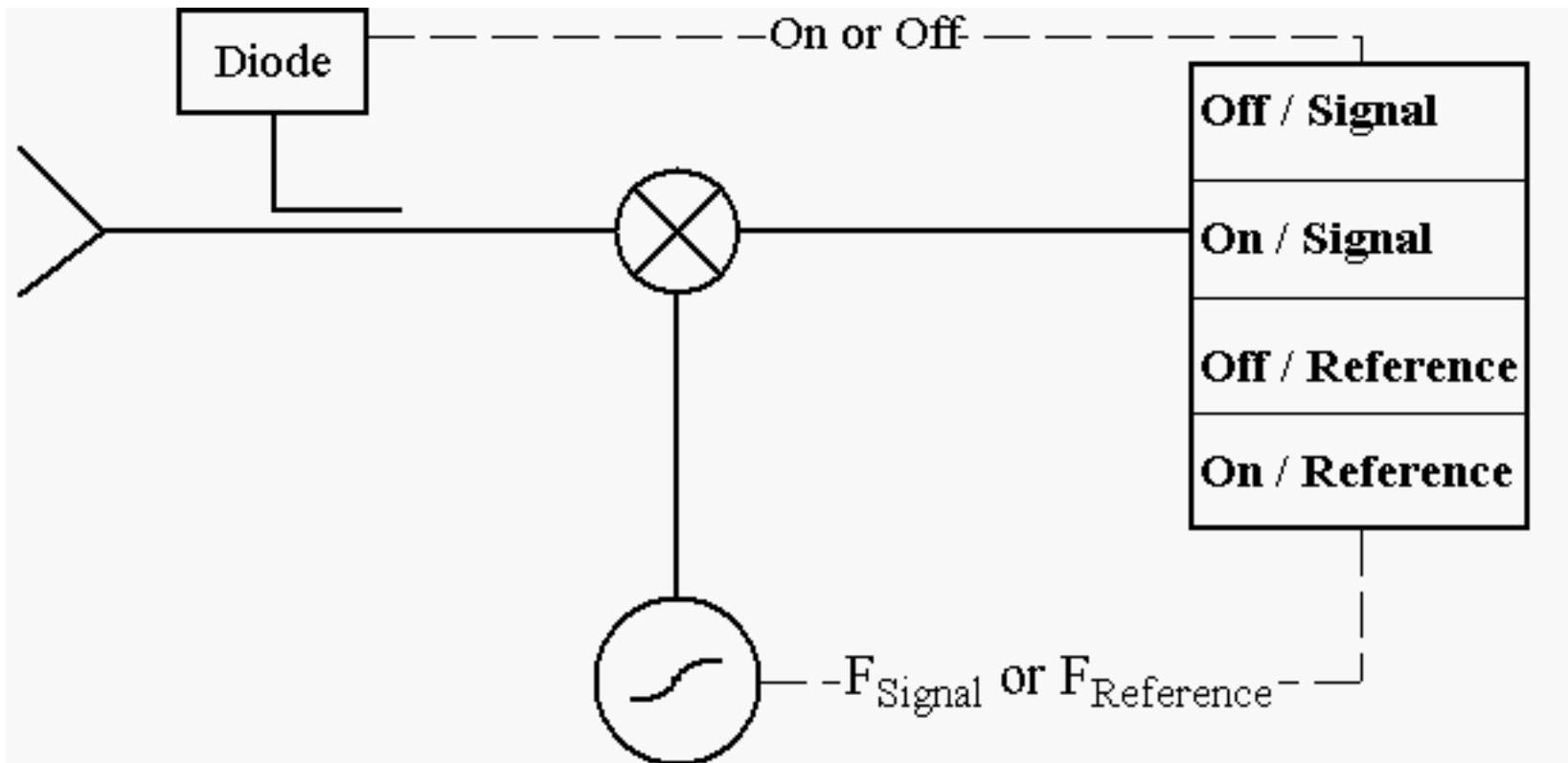
```
configLine = ""  
receiver = "Rcvr1_2"  
beam = "B1"  
obstype = "Spectroscopy"  
backend = "Spectrometer"  
nwin = 2  
restfreq = 1420.4058, 1665.0  
deltafreq = 0, 0  
bandwidth = 12.5  
swmode = "tp"  
swtype = "none"  
swper = 1.0  
swfreq = 0.0, 0.0  
tint = 30
```

```
vlow = 0  
vhigh = 0  
vframe = "lsrk"  
vdef = "Radio"  
noisecal = "lo"  
pol = "Linear"  
nchan = "low"  
spect.levels = 3  
""
```



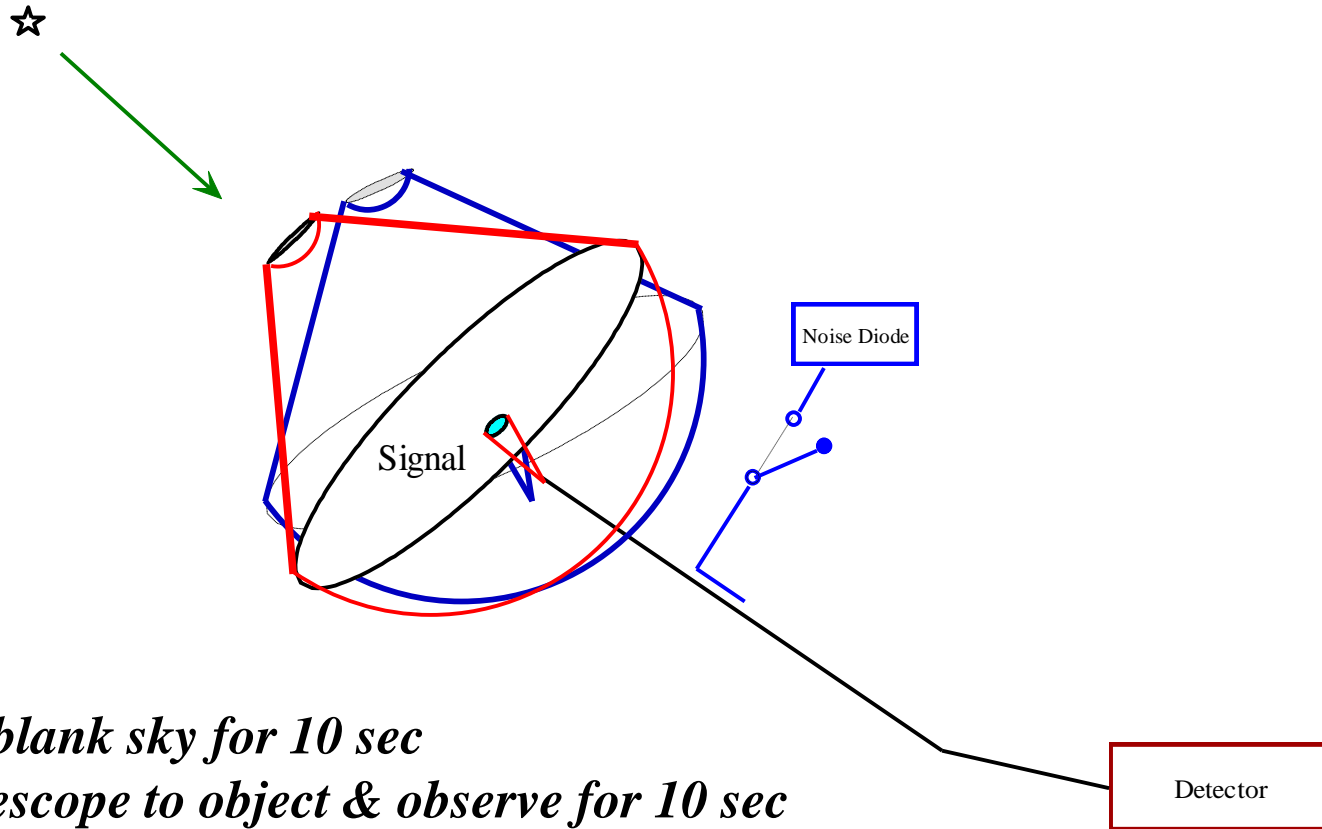


Model Receiver



Continuum - Point Sources

On-Off Observing

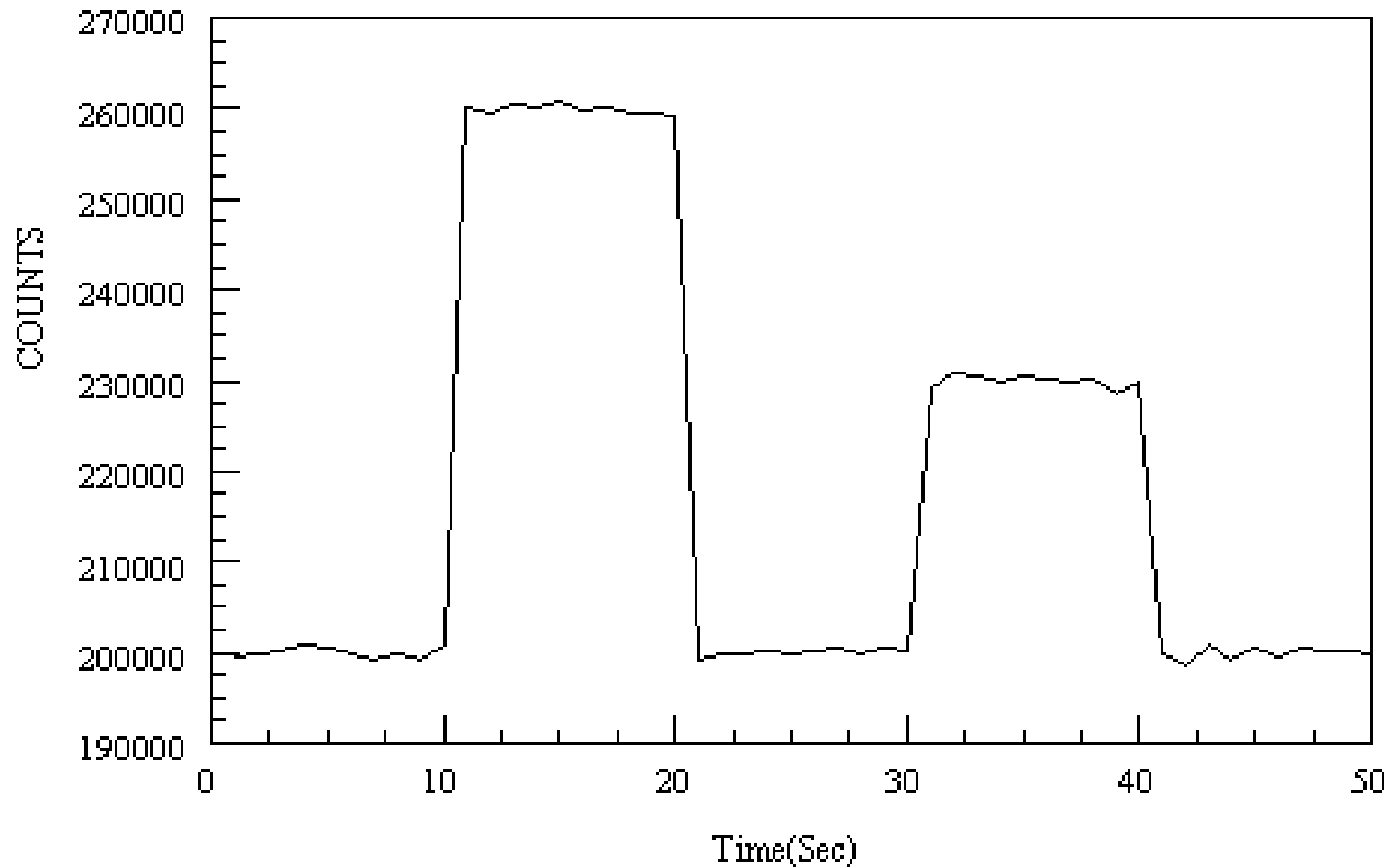


- *Observe blank sky for 10 sec*
- *Move telescope to object & observe for 10 sec*
- *Move to blank sky & observe for 10 sec*
- *Fire noise diode & observe for 10 sec*
- *Observe blank sky for 10 sec*



Continuum - Point Sources

On-Off Observing

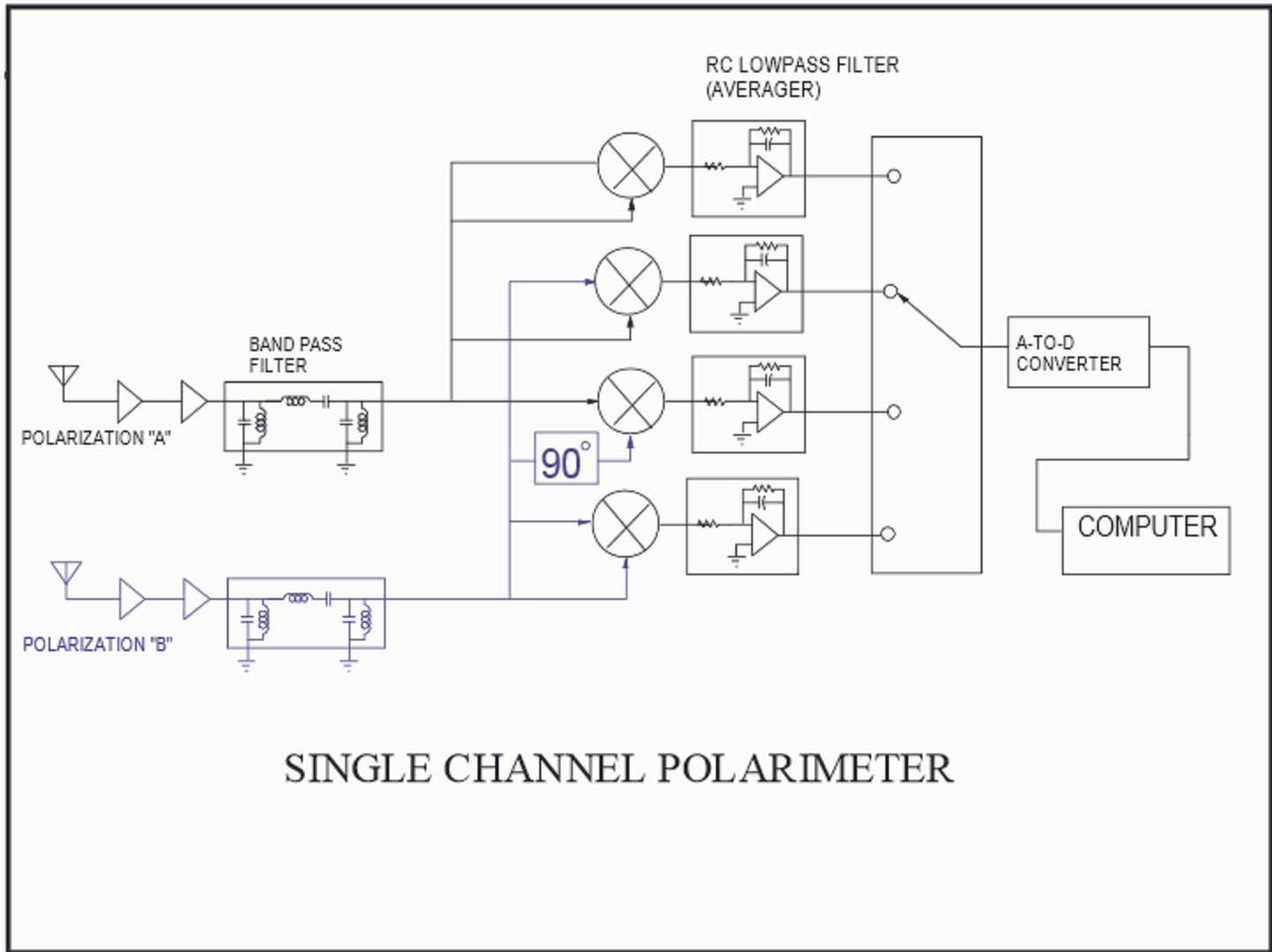


Continuum - Point Sources

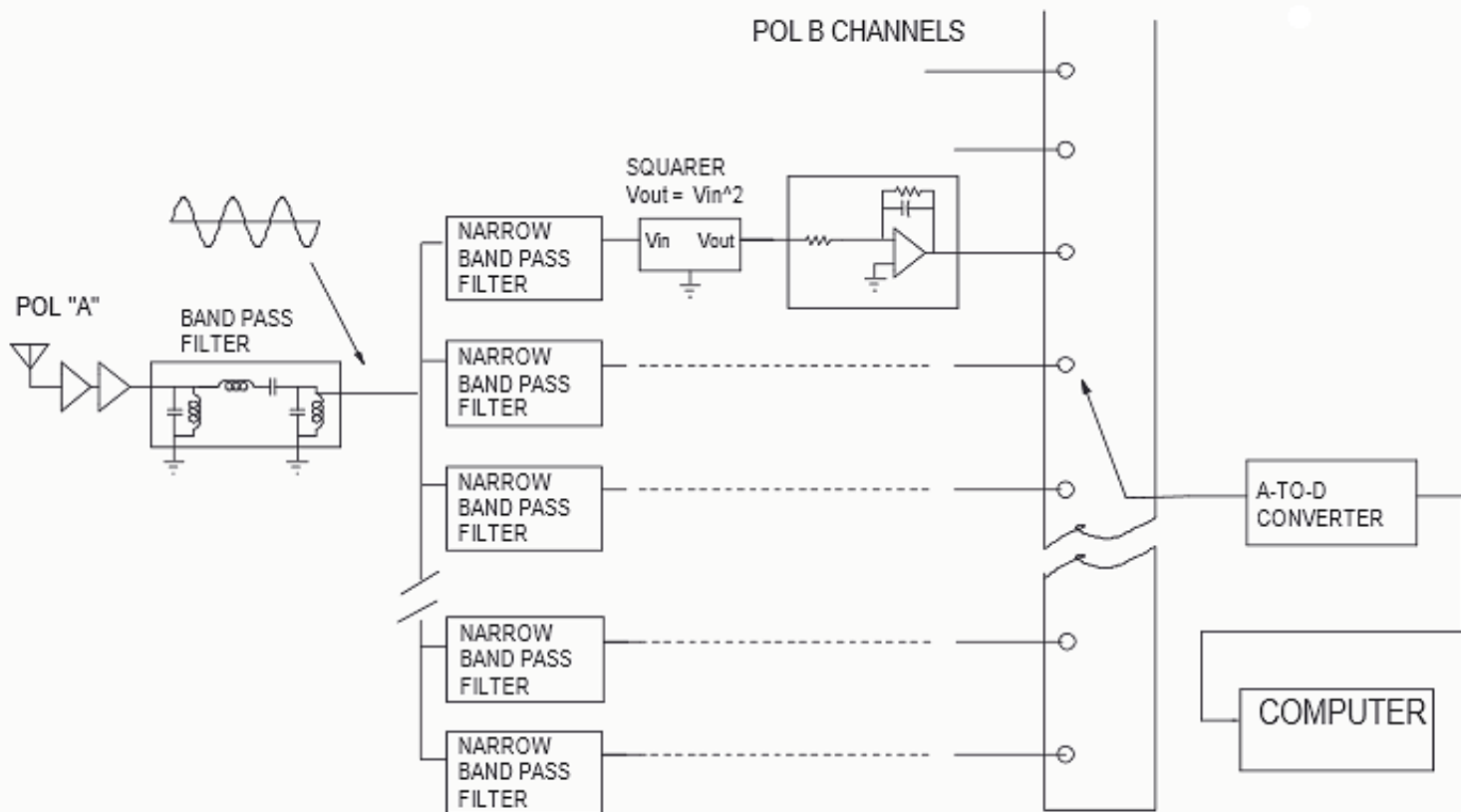
On-Off Observing

- Known:
- Equivalent temperature of noise diode or calibrator $(T_{\text{cal}}) = 3 \text{ K}$
- Bandwidth $(\Delta\nu) = 10 \text{ MHz}$
- Gain = 2 K / Jy
- Desired:
- Antenna temperature of the source (T_A)
- Flux density (S) of the source.
- System Temperature (T_s) when OFF the source
- Accuracy of antenna temperature (σ_{T_A})

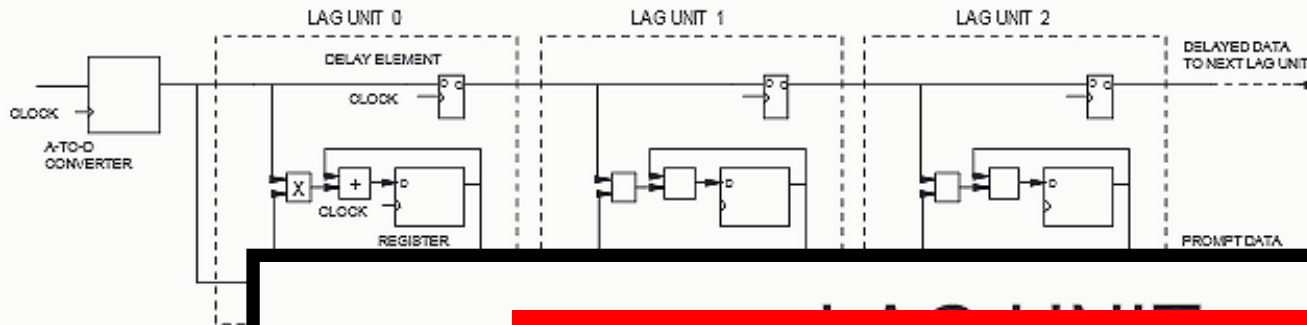




FILTER BANK SPECTROMETER



AUTOCORRELATOR



$$R(t) = \sum_{n=0}^{\text{Number Lags}} V(t) \cdot V(t + n \cdot \Delta t)$$

DELAYED DATA
(DELAYED IF S)

$$S(\text{Frequency}) = \text{DFT}(R(t))$$

$$\text{Bandwidth} = \frac{1}{2\Delta t} = \frac{\text{Sampling Frequency}}{2}$$

PROMPT DATA
(IF S)

$$\text{Resolution} = \frac{\text{Bandwidth}}{\text{Number of lags}}$$

DATA OUT

DATA OUT

